

## REFERENCES FOR TABLE XXIV

1. E. Grüneisen, in "Handbuch der Physik" (H. Geiger and K. Scheel, eds.), Vol. 10, p. 1. Springer, Berlin, 1962; NASA Natl. Aeron. Space Admin., Rept. (English Transl.) NASA-RE-2-18-59W.
2. J. J. Gilvarry, *J. Chem. Phys.* **23**, 1925 (1955).
3. M. W. Zemansky, "Heat and Thermodynamics," 4th ed. McGraw-Hill, New York, 1957.
4. M. H. Rice, R. G. McQueen, and J. M. Walsh, *Solid State Phys.* **6**, 1 (1958).
5. N. F. Mott and H. Jones, "Theory of the Properties of Metals and Alloys." Dover, New York, 1958.
6. D. B. Fraser and A. C. H. Hallett, *Proc. 7th Intern. Conf. Low Temp. Phys., Toronto, Ont., 1960*, p. 689. Univ. of Toronto Press, Toronto, Canada, 1961.
7. L. V. Al'tshuler, S. B. Kormer, A. A. Bakanova, and R. F. Trunin, *Zh. Eksperim. i Teor. Fiz.* **38**, 790 (1960); *Soviet Phys. JETP (English Transl.)* **11**, 573 (1960).
8. R. G. McQueen and S. P. Marsh, *J. Appl. Phys.* **31**, 1253 (1960).
9. W. J. Carter, *U.S. At. Energy Comm. Rept. LAMS-2640* (May, 1962).
10. K. A. Gschneidner, Jr., "Rare Earth Alloys." Van Nostrand, Princeton, New Jersey, 1961.
11. L. V. Al'tshuler, A. A. Bakanova, and R. F. Trunin, *Zh. Eksperim. i Teor. Fiz.* **42**, 91 (1962); *Soviet Phys. JETP (English Transl.)* **15**, 65 (1962).
12. V. T. Deshpande and D. B. Sirdeshmukh, *Acta Cryst.* **14**, 355 (1961).

seen, especially for the elements in the fourth, fifth, and sixth periods. The value for the alkali metals, which is close to the mean for all the metals, decreases as one moves along the period, with a minimum occurring at about the group IIIA elements. The value then gradually rises to a maximum at about the group IB metals, with a few minor oscillations between the group IIIA and IB metals. From the maximum at group IB,  $\gamma_G$  decreases almost steadily until the end of the row is reached.

The Grüneisen constants for the rare earths are shown in Fig. 33a, and it is seen that  $\gamma_G$  increases with increasing atomic number. The values for  $\gamma_G$  scatter considerably about the straight line drawn through the points.

The Grüneisen constant,  $\gamma_G$ , as calculated from the total heat capacity at constant volume,  $C_v$ , is always equal to, or less than, that calculated from  $C_p$ . The mean value of  $\gamma_G$  (from  $C_p$ ) is  $1.50 \pm 0.78$ . The error  $\pm 0.78$  corresponds to a percentage error of  $\pm 52.0\%$ . Comparing these calculated  $\gamma_G$  values with the literature values we find that 33 out of 51 of these values agree with one another within  $\pm 10\%$ . Most of the differences between the other 18 sets of values can be explained on the basis of revised thermal expansion and compressibility data. Such differences are to be expected, since many of the  $\gamma_G$  literature values are based on data over 30 years old. Thus, the agreement between literature values and those calculated herein seem to be reasonably good.

Grüneisen Constant  $\gamma_S$ . A comparison of  $\gamma_G$  (from  $C_p$ ) and  $\gamma_S$  was

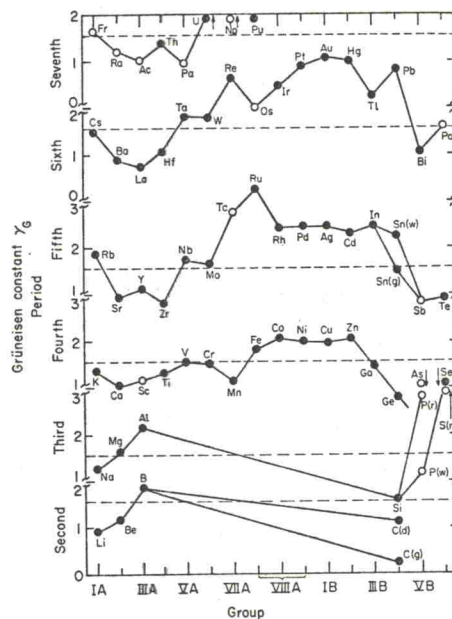


FIG. 32. Grüneisen constant,  $\gamma_G$ , as calculated from  $C_p$ , of all of the elements considered in this review with the exception of the rare-earth metals. The horizontal dashed line represents the mean values. Open points are estimated data.

made, thus extending analysis of Slater<sup>17</sup> and of Gilvarry<sup>117</sup> to all of the elements for which experimental data exist. Arbitrarily the following criteria were used in assessing the data: (1) if  $\gamma_S$  lies within  $\pm 25\%$  of  $\gamma_G$ , then the agreement was considered to be very good; if  $\gamma_S$  lies between  $\pm 25$  and  $\pm 50\%$  of  $\gamma_G$ , then the agreement was considered to be good; (3) if  $\gamma_S$  lies between  $\pm 50$  and  $\pm 80\%$ , the agreement was considered to be fair; and (4) if it was more than  $\pm 80\%$ , the agreement was considered to be poor. This scale is about the same as Gilvarry used.<sup>117</sup> It was found that for 18 elements the agreement was very good (B, Na, Al, K, Ca, Fe, Zn, Sr, Y, Nb, Cd, Sn(w), Cs, Pr, Ho, Er, Tm, and Yb); for 11 elements the agreement was good (Mg, V, Cr, Ni, Cu, Se, In, Nd, Sm, Au, and Tl); for 11 elements the agreement was fair (Li, Ti, Mo, Rh, Te, Ba, Dy, Hf, Pb, Th, and Pu); and for 20 elements the agreement was